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**(D1) Dust and Young Stars in Galaxies**

(D1.1.1) The UV slope of such a typical galaxy is  ;

(D1.1.2) See *Figure 1*;

(D1.1.3) The UV slope of CR7 is  ; [YES] or [NO]?  ;

(D1.2.1) See *Figure 2*; The best-fit Equation is:

;

(D1.2.2) The dispersion is  dex;

(D1.3.1)  $A_{1600}$  as a function of IRX:

;

(D1.3.2) See *Figure 3*; The Best-fit Equation is:

;

(D1.3.3) Expected UV slope of a dust-free galaxy:  ;

(D1.4.1) The IRX of CR7 is  ;

Here the value is a  limit (please fill with [upper] or [lower]);

(D1.4.2) Is the current observation deep enough ([YES] or [NO])?  .

**(D2) Compact Object in a Binary System**

(D2.1.1) The maximum acceleration is  km/s/day;

(D2.1.2) Estimate for the mass of the companion:   $M_{\odot}$ ;

(D2.2.1) See *Figure 4*;  $P_{orb} =$   ;  $K =$   ;

(D2.2.2)  $f(M_1, M_2) =$    $M_{\odot}$ ;

(D2.3.1)  $\sin i =$   ;  $R_1 =$    $R_{\odot}$ ;

$L_1 =$    $L_{\odot}$ ;  $M_1 =$    $M_{\odot}$ ;

(D2.3.2) The type of this star should be  ;

(D2.3.3) See *Figure 5*;

(D2.3.4) The possible mass of the unseen companion is   $M_{\odot}$ ;

It could be a  .